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1	Advances in the Estimation of High Spatio-Temporal Resolution Pan-African			
2	Top-Down Biomass Burning Emissions made using Geostationary Fire			
3	Radiative Power (FRP) and MAIAC Aerosol Optical Depth (AOD) Data			
4				
5	Authors:			
6	Hannah M. Nguyen <sup>a,c</sup> and Martin J. Wooster <sup>a,b</sup>			
7	<sup>a.</sup> King's College London, Department of Geography, Aldwych, London WC2B 4BG,			
8	UK.			
9	b. National Centre for Earth Observation, (NCEO), King's College London, Aldwych,			
10	London WC2B 4BG, UK.			
11	<sup>c.</sup> Centre for Cross-Disciplinary Approaches to Non-Equilibrium Systems (CANES),			
12	King's College London, Strand, London WC2R 2LS, UK.			
13				
14	Corresponding Author: Hannah M. Nguyen, hannah.nguyen@kcl.ac.uk			
15				
16	Abstract			
17	We provide major updates to the 'top down' Fire Radiative Energy Emissions' (FREM) approach			
18	to biomass burning emissions calculations, bypassing the estimation of fuel consumption that			
19	a major source of uncertainty in widely used 'bottom up' approaches. The FREM approach link			
20	satellite observations of fire radiative power (FRP) to emission rates of total particulate matte			
21	(TPM) via spatially varying smoke emissions coefficients (g.MJ <sup>-1</sup> ) – each derived from matchups			

22 of FRP and smoke plume aerosol optical depth (AOD). In the original FREMv1 approach, FRP 23 data came from the geostationary Meteosat satellite and AOD data from the 10 km spatial 24 resolution MODIS MOD04 aerosol product. However, the latter often performs guite poorly close 25 to biomass burning sources due to its large 10 km pixels, bias at high MODIS view zenith angles, 26 and saturation and/or removal of areas of high AOD - limitations introducing bias and uncertainty 27 into the final FREM-derived smoke emissions estimates. We address each of these issues 28 through a series of significant methodological and input data improvements, including exploitation 29 of the 1 km MODIS MAIAC AOD product that performs far better close to fire sources. We use 30 our FREMv2 methodology to generate a new pan-African fire emissions inventory for TPM and 31 the carbonaceous gases CO<sub>2</sub>, CO and CH<sub>4</sub>, and our annual mean TPM emissions are within 11% 32 of those of the MODIS-based FEER top-down approach, but significantly higher than those of 33 GFASv1.2 and GFEDv4.1s (by 114% and 69% respectively) - agreeing with independent 34 assessments that aerosol emissions of GFASv1.2 require upscaling by a factor of 2 to 3.4 to 35 deliver matching magnitudes between modelled and observed AODs. From our carbonaceous 36 emissions totals we map dry matter consumed (DMC) across Africa, and dividing this by the 37 FireCCISFD11 20 m burned area product we provide one of the first data-driven pan-African maps 38 of fuel consumption per unit area (kg.m<sup>-2</sup>) which in many areas is higher than in GFEDv4.1s. Our 39 estimates represent the highest spatio-temporal resolution biomass burning emissions data yet 40 available over Africa, and strongly advance the aim of a pan-tropical and mid-latitude inventory 41 based on FRP from the global geostationary satellite network (Meteosat, Meteosat IOD, GOES 42 and Himawari).

43

#### 45 1. INTRODUCTION

46 Biomass burning is amongst the largest contributors of gaseous and particulate emissions to the 47 atmosphere, generating a significant fraction of the global atmospheric load of black carbon (BC), 48 particulate matter (PM) and carbon monoxide (CO) (Andreae and Merlet, 2001; Forster et al., 49 2007; Reddington et al., 2016). Emissions impact regional and global air guality, weather and 50 climate variability (Crutzen and Andreae, 1990; Randerson et al., 2006; Westerling et al., 2006; 51 Schultz et al., 2008; Akagi et al., 2011). The highly dynamic spatio-temporal nature of landscape 52 fires makes their emissions challenging to quantify and satellite Earth Observation (EO) offers the 53 only means to do so over large spatial scales, especially where near real-time information is 54 required. However, despite advancements in the guality of EO data and in fire emission inventory 55 methodologies (Seiler and Crutzen, 1980; Flannigan and Vonder Haar, 1986; Pereira et al., 1999; 56 Justice et al., 2002; Wooster et al., 2005; Van Der Werf et al., 2006; Ichoku et al., 2008; Vermote 57 et al., 2009; Lehsten et al., 2009; Wiedinmyer et al., 2011; Kaiser et al., 2012; Ichoku and Ellison, 58 2014; Darmenov and da Silva, 2015; Mota and Wooster, 2018), large uncertainties and 59 discrepancies remain between the different inventories.

60 Most fire emission inventories use a 'bottom-up' approach, in which estimates of burned biomass 61 are generated from EO-derived metrics of burned area (BA), active fire counts and/or fire radiative 62 power (FRP). These burned biomass estimates are multiplied by biome-specific emission factors 63 (EFs) to relate each kilogram of burned dry matter to the amount of a trace gas or aerosol released 64 into the atmosphere. EFs are mostly derived from small scale laboratory or ground-based field 65 measurements (Andreae and Merlet, 2001; Akagi et al., 2011; Andreae, 2019), and more rarely 66 through airborne sampling of larger plumes (Abel et al., 2003; Lavrov et al., 2006; Quennehen et 67 al., 2012). 'Bottom up' emissions inventories include GFED (Van Der Werf et al., 2006, 2010, 68 2017), GFAS (Kaiser et al., 2012), FLAMBE (Reid et al., 2009) and FINN (Wiedinmyer et al., 69 2011). Biases and uncertainties present in these landscape fire inventories stem primarily from:

i) Limitations of the original satellite observations and the fire detection and
 characterisation algorithms applied to them to generate the EO-derived fire
 metrics. Compromises are generally made between spatial and temporal
 resolution, and algorithm errors of omission and commission impact the
 precision and accuracy of the EO-derived fire measures (Boschetti et al.,
 2004; Freeborn et al., 2009; Randerson et al., 2012).

Assumptions associated with estimating the fuel consumption per unit area
(kg.m<sup>-2</sup>) or any alternative scalar required to turn the EO-derived metric into
an estimate of burned dry matter (Kasischke and Penner, 2004; Reid et al.,
2009; Wooster et al., 2011; Kaiser et al., 2012)

80 iii) Limitations in the EFs used to convert between burned dry matter and the
81 final emissions of aerosols and trace gases (Van Leeuwen and Van Der Werf,
82 2011).

83

84 Addressing (i) above, advancements continue to be made to the EO-derived fire metrics extracted 85 from data collected by polar-orbiting sensors such as MODIS and VIIRS (e.g. Schroeder et al., 86 2014; Giglio et al., 2016; Zhang et al., 2017), and by geostationary sensors such as Meteosat 87 SEVIRI and Himawari-8 AHI (Wooster et al., 2015) and in the case of (iii), more detailed EFs are 88 regularly being proposed (e.g. Akagi et al., 2011; Huijnen et al., 2016; Andreae, 2019). However, 89 arguably less research has focused on (ii), namely, how to improve estimates of burned dry matter 90 derived from burned area, active fire (AF) detection or FRP data products. Despite this being 91 considered to be the step introducing probably the greatest uncertainties in 'bottom-up' 92 approaches (Reid et al., 2009; Ichoku and Ellison, 2014; Mota and Wooster, 2018). Partly for this 93 reason, fully 'top-down' methodologies such as those of Ichoku and Ellison (2014) and Mota and 94 Wooster (2018) have taken to deriving landscape fire emissions estimates directly from EO-

95 derived FRP measures, thereby bypassing the fuel consumption estimation step altogether and 96 reducing the number of assumptions required during the fire emissions calculation. In these 97 approaches, a biome-dependent scalar (a smoke emission coefficient;  $C_e$  in g.MJ<sup>-1</sup>) captures the 98 relationship between the FRP of fires and their associated total particulate matter (TPM) 99 emissions. These coefficients are derived from a series of matchup fires where FRP data and 100 satellite aerosol optical depth (AOD) observations are available in the biome of interest. Once this 101 scalar is determined, the need to calculate fuel consumption is removed when deriving further 102 smoke emissions estimates from the FRP data of observed fires. Whilst these 'top-down' 103 approaches successfully bypass the fuel consumption step, in both the Fire Energetics and 104 Emissions Research (FEER; Ichoku and Ellison, 2014) and Fire Radiative Energy Emissions 105 (FREM; Mota and Wooster, 2018) approach, the coarse 10 km spatial resolution of the MODIS 106 AOD product used to derive in-plume TPM, and performance issues related to this AOD product 107 in thick-smoke affected environments, can introduce significant problems when deriving the 108 smoke emission coefficients. To address this problem, we here present a series of improvements 109 to the FREMv1 methodology of Mota and Wooster (2018) and use this (FREMv2) method to 110 produce a new Meteosat SEVIRI FRE-based fire emissions inventory for the whole of Africa. A 111 series of methodological evolutions are presented, key of which is the exploitation of a far higher 112 spatial resolution (1 km) MODIS AOD product (Lyapustin et al., 2018) that offers improved 113 performance in heavily smoke impacted environments.

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#### 115 2. FIRE EMISSION INVENTORY METHODOLOGIES

### 116 2.1 'Bottom-up' Total Fuel Consumption (*F<sub>c</sub>*) Based Methodologies

117 Most widely used 'bottom-up' fire emission inventories currently take their emission factors (EFs 118 in g.kg<sup>-1</sup>) from those collated by Andreae and Merlet (2001), although recent updates are reported 119 in Andreae (2019). Therefore, inter-inventory variations in both total fire emissions magnitude and 120 spatio-temporal distribution originate primarily not from the EFs used, but in the estimation of total 121 fuel consumption ( $F_c$ ), often called dry matter consumed (DMC). The uncertainty in DMC 122 estimates primary stems from limitations (i) and (ii) introduced in Section 1. These are 123 performance issues in the EO data products (typically BA or FRP) used to characterise the 124 biomass burning, and the assumptions made when estimating DMC from these (e.g. Wooster et 125 al., 2011; Knorr et al., 2012; Larkin et al., 2014).

126 In the Burnt Area (BA) approach to estimating total fuel consumption (Seiler and Crutzen, 1980), 127 DMC is estimated via multiplication of BA [m<sup>2</sup>], fuel load [FL; kg.m<sup>-2</sup>] and combustion 128 completeness [CC: unitless]. BA measures are typically taken from the 500 m spatial resolution 129 MODIS MCD64A1 product (Giglio et al., 2018), with the parameters of fuel load and combustion 130 completeness either taken from literature values (e.g. Reid et al., 2005) or the outputs from 131 vegetation growth models (GFED; van der Werf et al., 2006). In addition to uncertainties in the 132 fuel load and combustion completeness parameters (Reid et al., 2009; Vermote et al., 2009; Van 133 Leeuwen and Van Der Werf, 2011), many small fires appear to remain unmapped in MCD64A1 134 (Tsela et al., 2014; Hawbaker et al., 2017; Roteta et al., 2019). A widely used BA-based fire 135 emissions inventory is the Global Fire Emissions Database (GFED; Van Der Werf et al., 2006), 136 providing emissions at 0.25° grid cell resolution. The latest version (GFEDv4.1s; Van Der Werf et 137 al., 2017) uses MODIS active fire (AF) detections to 'boost' the BA values in an attempt to 138 compensate for small fires remaining undetected by MCD64A1 (Randerson et al., 2012). Whilst MCD64A1 BA totals clearly need to be raised in some way, this 'boosting' process can introduce
biases in BA in regions that are, for example, dominated by interspersed agricultural and urban
areas (Zhang et al., 2018).

142 In the FRP-based approach to estimating dry matter consumed, FRP observations are temporally 143 integrated to estimate Fire Radiative Energy (FRE), which in small-scale fires has been shown to 144 directly relate to DMC (Wooster et al., 2005; Freeborn et al., 2008; Ichoku et al., 2008). A 145 conversion factor (g.MJ<sup>-1</sup>), typically derived from these small-scale fires (e.g. that of Wooster et 146 al. (2005)) is then used to convert satellite-based FRE measures into DMC estimates (Roberts et 147 al., 2005, 2011). This removes the need for assumptions about the poorly constrained parameters 148 of fuel load and combustion completeness that are required by methods using BA data – but relies 149 on the fact that the conversion factor derived from small scale fire experiments remains valid at 150 the scale of landscape fires observed from satellites. This may be an unrealistic assumption, 151 particularly in certain landscapes (Mota and Wooster, 2018). For example, tree canopy cover may 152 intercept a significant percentage of the fire-emitted FRP in more forested regions, leading to 153 lower spaceborne FRP measurements (Freeborn et al., 2009; Mota and Wooster, 2018; Roberts 154 et al., 2018a). A positive benefit of satellite AF products are that they are sensitive to small fires 155 covering as little as 10<sup>-4</sup> of the pixel area (Roberts et al., 2005), hence the contribution of relatively 156 small fires is typically included in FRP-based fire products, provided they are burning at the time 157 of the satellite observation. This compares to most satellite BA products, where generally > 20% 158 of the pixel area must be fire affected for a burn scar to be reliably detectable (Giglio et al., 2006. 159 2009), but with the advantage that the burned area remains detectable and measurable for some 160 time after the fire has ceased.

For providing near-real time fire emissions estimates, BA-based approaches are, in any case,usually inappropriate since the required burned area, FL and CC datasets are typically only

163 available with significant time-lag. This is one further reason why FRP-based approaches are 164 becoming more common, since satellite FRP data of a fire are typically available within minutes 165 to hours depending on the observation system. A widely used FRP-based fire emissions 166 inventory is the Global Fire Assimilation System (GFAS; Kaiser et al., 2012). GFAS provides daily 167 data at 0.1° globally as part of the Copernicus Atmospheric Monitoring Service (CAMS: 168 https://atmosphere.copernicus.eu/). In GFAS, weighted mean FRP values from MODIS 169 MOD/MYD14 products (Justice et al., 2002; Giglio et al., 2003) are used to derive daily FRE 170 estimates (assuming either a flat FRP emissions profile or a standard fire diurnal cycle), which 171 are then multiplied by biome-specific conversion factors (in kg.MJ<sup>-1</sup>) relating FRE to DMC (Kaiser 172 et al., 2012). These GFAS-specific conversion factors are based on prior comparisons between 173 the MODIS-based GFASv1.0 FRE estimates and the BA-derived DMC totals of GFEDv3.1 (Van 174 Der Werf et al., 2006, 2010), which themselves come from an adaption of the original Seiler and 175 Crutzen (1980) approach (Heil et al., 2010). In this way, GFAS fire emissions estimates are not 176 only influenced by GFEDs biases and uncertainties in fuel load, combustion completeness and 177 unmapped small burned areas (Reid et al., 2009; Vermote et al., 2009), but are also directly 178 dependent on the GFED DMC estimates themselves (Kaiser et al., 2012). The operational GFAS 179 also uses only polar-orbiting satellite FRP data, currently that from MODIS, somewhat limiting its 180 temporal resolution and the accuracy of any FRP-to-FRE conversion.

181 2.2 FRP Datasets for Emissions Estimation

The FRP measures derived from the polar-orbiting (MODIS) sensors used within GFAS and elsewhere fail to capture the full diurnal cycle of a fire, and must typically be interpolated between observations or used to scale an assumed diurnal cycle in order to estimate FRE (Freeborn et al., 2009; Vermote et al., 2009). It is however possible to use the very high temporal resolution FRP data available from geostationary satellites to provide almost continuous FRP observations, and

187 these can then be accurately and easily integrated to calculate FRE without any assumptions on 188 the shape of the fire diurnal cycle (e.g. Wooster et al., 2015). The main limitation of geostationary 189 FRP datasets is that the minimum FRP detection limit, below which actively burning fires remain 190 undetected, is higher than for most polar-orbiting sensors due to the larger geostationary pixel 191 areas (Roberts et al., 2005, 2015). However, the detectable AFs in geostationary products still 192 remain significantly smaller in terms of pixel area coverage (e.g. down to perhaps 0.01% of the 193 pixel) than the minimum burned area detectable in the MODIS BA products. Another 194 disadvantage of geostationary active fire data is that geostationary pixel areas grow markedly at 195 locations very far from nadir (>  $40^{\circ}$  view zenith angle), so high latitude regions are generally less 196 well suited to geostationary FRP analysis. A recent evaluation of geostationary active fire data 197 was carried out by Hall et al. (2019) who compared AF detections within the Meteosat SEVIRI 198 FRP-PIXEL product used herein (see Section 4) to those identified in 30 m spatial resolution 199 Landsat-8 Operational Land Imager (OLI) data (Hall et al., 2019). Results for the AF error of 200 commission showed this to be 8% for the SEVIRI product compared to OLI, a false alarm rate 201 very similar to that of the widely used MODIS AF products (Giglio et al., 2013; Schroeder et al., 202 2014). AF errors of omission for the SEVIRI product were however 98% compared to OLI, but the 203 30 m OLI pixel size enables fires covering just a few square meters to be detected, compared to 204 the roughly one thousand square meters required for detection with SEVIRI assuming an AF 205 detection limit of 0.01% of the pixel area. Thus, a large number of very small fires can be detected 206 by OLI that remain essentially impossible to detect by SEVIRI. Comparison between the SEVIRI 207 FRP-PIXEL AF product and 1 km MODIS AF data have been commonly performed (Roberts and 208 Wooster, 2008; Roberts et al., 2015) and the SEVIRI AF error of omission rate in such 209 comparisons is far lower than when OLI is used as the reference dataset (Roberts et al., 2015). 210 Indeed, Wooster et al. (2015) indicate that over the lifetimes of most African fires, active fires 211 detected by MODIS

become detectable by SEVIRI at some point in their diurnal cycle. The Hall et al. (2019) study also highlights that, though at any given time the higher spatial resolution polar orbiter-based FRP products are likely to detect more fires than geostationary based products due to their ability to detect lower FRP fires, the increased temporal resolution provided from geostationary orbit typically results in far more AF detections overall, during a 24 hour period.

218 2.3 Top-down FRP-based Methodologies

219 The FREMv1 approach to fire emissions estimation (Mota and Wooster, 2018) is an FRP-based 220 calculation, classed as a 'top-down' method since it only uses satellite observations (specifically 221 geostationary FRP and polar orbiting-derived Aerosol Optical Depth [AOD] data). The method 222 avoids problems inherent in the intermediate DMC estimation step of the 'bottom-up' approaches 223 (Section 2.1), though fuel consumption in terms of total DMC or its combustion rate can still be 224 calculated as a final output (see Mota and Wooster, 2018). Whilst the GFAS system (Kaiser et 225 al., 2012) is also based on FRP data, it is not fully top down as it relies on conversion coefficients 226 between FRP and DMC that come from BA-based approaches which themselves rely on model-227 based fuel load and combustion completeness variables (Van Der Werf et al., 2006, 2010). The 228 FREM approach is somewhat similar to the fire emissions estimation approach introduced by 229 Ichoku and Kaufman (2005) in that it directly links FRP data to emissions of total particulate 230 matter, albeit FREM uses geostationary rather than polar-orbiting FRP data for the reasons 231 discussed in Section 2.2. Ichoku and Kaufman (2005) and follow-up work by Ichoku and Ellison 232 (2014) took both their FRP and AOD data from the polar-orbiting MODIS sensor, and used these 233 datasets to deliver the Fire Energetics and Emissions Research (FEER) smoke emission product. 234 In the FEER and FREM approaches to fire emissions estimation, as well as in the other top-down 235 methods of Lu et al. (2019) and Darmenov and da Silva (2015), the intermediate step of estimating 10

236 DMC (kg) or dry matter combustion rate (kg.s<sup>-1</sup>) is bypassed by deriving a 'smoke emission 237 coefficient'  $[C_e]$  describing the relationship between the thermal energy a fire radiates (i.e. the FRE 238 in MJ) and the mass of total particulate matter (TPM in kg or g) it emits, or between the rates of 239 these two (i.e. the FRP in MW and the TPM emission rate in  $g.s^{-1}$ ).  $C_e$  has units of  $g.MJ^{-1}$  or  $g.s^{-1}$ 240 <sup>1</sup>.MW<sup>-1</sup> respectively in each case, and is itself typically derived from a set of matchup fires for 241 which good observations of both variables exist. To derive  $C_{e}$ , each matchup fire has its TPM 242 estimated using satellite observations of AOD, or from modelled AOD in the case of Darmenov 243 and da Silva (2015). Once  $C_e$  is determined using these matchup fires, it can be applied to the 244 FRE or FRP data of all fires to estimate their TPM emissions or emissions rate as well. FEER and 245 Lu et al. (2019) use MODIS FRP data as opposed to FREM's geostationary FRP data, and 246 therefore can be affected by MODIS's over-or-under estimation of FRE due to its limited temporal 247 resolution as discussed in Section 2.2 (Kaiser et al., 2012; Andela et al., 2015). The "bow tie effect" caused by the sensor's design and scanning geometry can also affect the estimation of 248 249 FRE from FRP measures (Freeborn et al., 2011; Wiedinmyer et al., 2011). Proposed methods to 250 address the estimation of FRE from MODIS FRP assume that the MODIS Aqua early-afternoon 251 overpass roughly coincides with, and captures, the peak of daily fire activity, and can therefore 252 be used to parameterise a diurnal cycle that provides interpolation-based higher temporal 253 frequency estimates of FRP (Ellicott et al., 2009; Vermote et al., 2009). However, due to the 254 considerable spatial variability seen in both fire diurnal cycles (Giglio, 2007) and the local MODIS 255 Aqua overpass time, the time difference between the peak of fire activity and the MODIS overpass 256 varies both geographically and daily, as does the fraction of total daily fire pixel counts occurring 257 at the MODIS overpass (Mota and Wooster, 2018). In cases where the latter is particularly low, 258 polar-orbiting based calculations of daily FRE are susceptible to a low bias, but when the MODIS 259 overpass coincides closely with peak fire activity this is less pronounced. Such spatial variations 260 can lead to artificial differences in the derived daily FRE (Mota and Wooster, 2018), and this is

261 one justification for the use of geostationary-derived FRP data in the FREM approach. 262 Furthermore, in the FEER approach of Ichoku and Ellison (2014) and in the Lu et al. (2019) 263 methodology, the AOD and FRP observations of the matchup fires used to derive the  $C_{\rho}$ 264 coefficients are based on MODIS observations of the plume and actively burning fire that are 265 acquired at exactly the same time. This inherently means that the TPM contained within the 266 plume (which has been emitted since the fire commenced up to the time of the MODIS overpass) 267 is being related to the FRP the fire is releasing at the moment of the overpass. This is a potential 268 disadvantage, because earlier on in the fire lifetime – during which most of the TPM was actually 269 released - the fire might have had a guite different FRP to that at the time of the MODIS overpass. 270 The FREM approach avoids this problem by using geostationary FRP observations and 271 integrating these over time to calculate the FRE released by the fire from the time it started until 272 the time of the MODIS overpass providing the AOD-based TPM estimate. Furthermore, by 273 subsequently applying the  $C_e$  coefficients to all the geostationary FRP data, not just the matchup 274 fires, very high temporal resolution fire emissions estimates can be generated, which can be 275 important for (i) capturing sub-daily variability in fire and smoke emissions (Roberts et al., 2015), 276 and (ii) delivering the most accurate modelling of atmospheric pollutant dispersion and peak air 277 pollutant concentrations in large fire events (Baldassarre et al., 2015).

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### 280 2.4 Further limitations in current top-down methodologies

In all previous top-down methodologies, biome dependant (e.g. FREMv1) and/or geographically
dependant (e.g. FEER) smoke emission coefficients have been derived using satellite-based FRP
observations and the 10 km AOD (MOD/MYD04) MODIS Level 2 product from Terra (~10:30
overpass) and Aqua (~ 13:30 overpass) (Ichoku and Kaufman, 2005; Ichoku and Ellison, 2014;

285 Mota and Wooster, 2018; Lu et al., 2019). The outputs of two different AOD retrieval algorithms
286 are available in this MODIS product;

i) Dark Target (DT): designed to work over dense vegetation.

288 ii) Deep Blue (DB): developed for desert surfaces but updated to apply to289 most cloud free land.

Output from Dark Target was used by Ichoku and Kaufman (2005) and Ichoku and Ellison (2014),
whereas Mota and Wooster (2018) and Lu et al. (2019) use Deep Blue (introduced in Collection
6) since it showed better agreement with AREONET AOD observations across southern Africa
(Sayer et al., 2014).

294 The MODIS 10 km AOD (MxD04) product and its updates (Tanre et al., 1997; Hsu et al., 2004, 295 2006, 2013; Remer et al., 2005; Levy et al., 2013) has been produced for more than 15 years, 296 are widely used within the air quality community (Remer et al., 2013), and have been extensively 297 characterised and validated (Holben et al., 2001; Chu et al., 2002; Ichoku et al., 2002, 2003; Sayer 298 et al., 2013; Livingston et al., 2014; Xiao et al., 2016; Jethva et al., 2019). However, their use in 299 top-down fire emissions inventories introduce some key uncertainties and limitations, including 300 those associated with its relatively low 10 km spatial resolution (Lyapustin et al., 2011; Raffuse et 301 al., 2013; Remer et al., 2013; He et al., 2017), the cloud mask implemented in both the DT and 302 DB versions (Levy et al., 2013; Raffuse et al., 2013; Remer et al., 2013; Livingston et al., 2014), 303 and MODIS view-angle effects (Saver et al., 2015). The more recent 3 km MODIS AOD product 304 which applies the Dark Target retrieval algorithm (Remer et al., 2013) also suffers from distortions 305 at wide swath angles and erroneous cloud masking over thick smoke.

AOD retrievals in the MxD04 product are produced at 10 km to improve the signal-to-noise ratio of the input reflectance data (Tanre et al., 1997). In the Deep Blue retrieval, henceforth referred to as MxD04\_DB, the original 1 km pixels are arranged into  $10 \times 10$  pixel blocks and screened to

309 remove those affected by cloud. The remaining pixels in each 10 km block are then used to deliver 310 the AOD retrieval (Hsu et al., 2013), and the number of 1 km pixels used accompany the final 10 311 km AOD estimates contained within the product. When targeting smoke plumes, our 312 investigations and those of other studies (e.g. Levy et al., 2013; Livingston et al., 2014) show that 313 the conservative cloud mask used by both the DT and DB algorithms result in cases of (i) the 10 314 km AOD estimates at the locations of the thickest smoke being completely masked out, and (ii) 315 the 10 km AOD estimates in areas of thick smoke being retrieved from only a fraction of the 1 km 316 pixels present within the 10 km  $\times$  10 km box. Figure 1 demonstrates an example of a smoke plume 317 where some MxD04 DB pixels use as little as 40% of the original one hundred MODIS 1 km 318 pixels to retrieve a single 10 km AOD value, though at wide swath other distortions occur and 319 pixels are in fact not 10 km<sup>2</sup>. A comparison with the 500 m MODIS true colour image for the same 320 area (Fig. 1c) demonstrates how the low spatial resolution of the MxD04 BD product introduces 321 further uncertainty into the determination of the plume boundary, particularly when compared with 322 the alternative 1 km MAIAC AOD product (Fig. 1d) discussed in Section 3.



324 (1.5 columns) Figure 1. Example plume from a fire burning north of the Save River (Mozambigue), 325 imaged on the morning of 8<sup>th</sup> October 2015 at 11:15 UTC via the Aqua satellites MODIS sensor 326 (at a VZA 40.6°). (a) 10 km MxD04 DB AOD product; (b) MxD04 DB field showing the number 327 of 1 km reflectance pixels (out of 100) used to retrieve each 10 km AOD pixel value; (c) 500 m 328 MODIS Corrected Reflectance (True Colour) image; and (d) 1 km MCD19 MAIAC AOD product 329 derived from the same MODIS imagery shown in (c). The colour scale shown in (d) is also relevant 330 for (a). The plume is far more easily distinguished in the 1 km than the 10 km AOD product and 331 better matched to the smoke spatial distribution shown in the MODIS true colour image of (c). 332 Unlike the 1 km AOD product of (d), the 10 km MxD04\_BD AOD data of (a) rather poorly defines 333 the plume bounds and some pixels in this product are heavily impacted by the cloud mask which 334 removes AOD pixels over the thickest smoke (b). Some erroneous masking does occur in the 1 335 km product of (d), shown as the black pixels, but this is minimal and addressed via the 336 interpolation described in Section 4.3.

337 The extreme masking of smoke affected 1 km pixels as cloud in the MxD04 DB algorithm, as 338 demonstrated in Figure 1, introduces clear uncertainty and probably bias into any estimates of 339 total particulate matter (TPM) derived from the 10 km MxD04 AOD observations. Since the 340 excluded pixels are mainly located over the thickest smoke, their exclusion is likely to significantly 341 affect the final retrieved 10 km AOD, and therefore the TPM measure. Additionally, the complete 342 masking of 10 km pixels in some cases limits the number of plumes that can be identified and 343 used in deriving the FREM smoke emission coefficients. Both Mota and Wooster (2018) and Lu 344 et al. (2019) take measures to try and minimise the effect of completely masked MxD04 BD 345 pixels. Mota and Wooster (2018) keep MxD04 BD pixels with quality mark  $\geq$  2, resulting in 346 MxD04\_BD pixels using as little as 40% of the native 1 km pixels being retained (e.g. in Fig. 1b), 347 and they also excluded plumes comprising any completely masked 10 km pixels. Lu et al. (2019) 348 use a nearest neighbour method to fill gaps in the data over plumes, which are caused by 349 erroneously 'cloud masked' AOD pixels.

350 Another source of uncertainty introduced when using either the MODIS DT or DB MxD04 product 351 comes from the MODIS 'bowtie' effect that results from the MODIS design and scan geometry, 352 along with Earth's curvature (Wolfe et al., 1998). Above view zenith angles (VZA) of ~ 20°, two 353 key distortions occur with respect to MxD04 BD - (i) growth of the MODIS 10 km AOD 'pixel' 354 area from about 10 x 10 km at nadir to around 20 x 40 km at the scan edge, and (ii) the overlap 355 of successive scans towards the scan edge meaning features are duplicated in adjacent pixels 356 (Wolfe et al., 1998). Both result in an AOD data product rather dependent on the location of the 357 AOD pixels within the MODIS swath, and there are indications that this may influence the statistics 358 of the AOD retrievals towards the scan edges (Sayer et al., 2015), potentially introducing bias into 359 top-down methodologies using the MxD04 AOD products without excluding high-VZA 360 observations during derivation of their  $C_e$  values.

361 A further significant limitation of the MxD04 AOD products relate to their fundamentally low 10 km 362 (at nadir) spatial resolution (see Figure 1), which is unable to resolve smoke plumes from many 363 smaller fires, or from fires not sufficiently isolated from other aerosol sources (primarily other 364 nearby fires). In both cases, differences between the in-plume AOD and the background AOD 365 (i.e. the AOD anomaly) may not be significant enough to define the AOD pixels that represent the plume. This both places a limit on the minimum fire size used to derive Ce or Cbiome values and 366 367 makes the sampling of fires over certain areas during periods of peak fire activity difficult, as their 368 plumes often merge together (Mota and Wooster, 2018).

369 Separate from the MxD04 AOD product issues, differences in the method chosen to calculate the 370 value of an AOD background exist between top-down methodologies. In all top-down inventories 371 discussed herein, the equations proposed by Ichoku and Kaufman (2005) are used to convert 372 AOD to the emitted TPM of an individual fire. The fire emitted AOD for a given smoke plume is 373 defined as the summed total AOD of the smoke plume above the AOD background (for full details 374 see Ichoku and Kaufman, 2005). It is clear then, that the choice of AOD background value used in these calculations impacts the final TPM estimate, and therefore the Ce or Cbiome values 375 376 derived from matchups between these values and the corresponding fire FRE. Ichoku and 377 Kaufman (2005) and Ichoku and Ellison (2014) select background AOD values based on pixels 378 immediately up-wind of the smoke plume, whereas Mota and Wooster (2018) take background 379 AOD values to be the 20<sup>th</sup> percentile AOD of all values within a set distance or area (e.g. 500 km<sup>2</sup>) 380 surrounding a fire. The background AOD values estimated from this large-scale averaging may 381 not be fully representative of the true background into which a plume is being emitted, and so 382 may also negatively impact the derived  $C_{biome}$  values in FREMv1.

The above problematic characteristics of the MxD04 products, and the FREMv1 and FEER methodologies, contribute to uncertainty and possibly bias in the smoke emission coefficients

derived, and thus to the resultant smoke emissions estimates. We here focus on developmentsto FREMv1 to try to mitigate these impacts.

387

#### 388 *3.* DEVELOPMENTS TO THE FREM APPROACH

389 The methodology proposed here builds on the original FREMv1 approach of Mota and Wooster 390 (2018) to address the limitations described in Section 2.3. A key advance is use of an alternative 391 MODIS AOD product that offers substantial advantages over the MxD04 products when used in 392 smoke-affected areas. This alternative AOD product is based on the Multiangle Implementation 393 of Atmospheric Correction (MAIAC) algorithm, developed to retrieve surface bidirectional 394 reflectance factor (SBRF), internal cloud mask and AOD over land (Lyapustin et al., 2011). The 395 MAIAC AOD product provides combined Aqua and Terra AOD retrievals at a 1 km resolution over 396 both dark and bright surfaces, and has been shown to improve the resolvability of atmospheric 397 smoke and dust features compared to the 10 km MxD04 product (Emili et al., 2011; Lyapustin et 398 al., 2011, 2012; Jethva et al., 2019; Mhawish et al., 2019). At AERONET station locations, both 399 Emili et al. (2011) and Jethva et al. (2019) show that the MAIAC product provides more than 400 double the number of AOD retrievals compared to MxD04\_DB, due to its higher spatial resolution 401 and improved cloud mask. In fact, the MAIAC algorithm explicitly includes a 'smoke test' to 402 discriminate biomass burning smoke from clouds (Lyapustin et al., 2012). Other evaluations have 403 compared the MAIAC product to the Visible and Infrared Imaging Radiometer Suite (VIIRS) 750 404 m spatial resolution AOD product, to AERONET measurements, and to surface measurements 405 of particulate matter (Hu et al., 2014; Arvani et al., 2016; Martins et al., 2017; Superczynski et al., 406 2017). They have shown its improved coverage compared to the standard VIIRS AOD product 407 and its good agreement with ground-based AOD and particulate matter measures.

408 In addition to its higher spatial resolution and improved cloud mask, some key features of the 409 MAIAC product address issues related to the dependence of AOD retrievals on VZA in the MxD04 410 product. These include the gridding of L1B MODIS bands to 1 km resolution prior to AOD retrieval 411 using an area-weighted method (Wolfe et al., 1998), and the calculation of surface BRF using a 412 dynamic spectral regression coefficient (SRC) (Lyapustin et al., 2011). The former results in an 413 improved representation of any given 1 km grid cell by appropriately weighing the contribution of 414 observations falling within that cell, and this is especially important at the swath edge where the 415 MODIS pixel area is up to eight times larger than at nadir. The dynamic SRCs, are calculated 416 from a time series analysis of previous AOD retrievals for each 1 km pixel. Therefore, when VZA 417 are well sampled in the preceding retrieval times series (multiple cloud free observations per 418 pixel), SRC values represent well the angular component of surface BRF (full details can be found 419 in Lyapustin et al. (2011)). The final MAIAC AOD product is reported on a 1 km grid in the MODIS 420 sinusoidal projection (Lyapustin et al., 2018) and Mahawish et al. (2019) show VZA-dependant 421 bias to be the lowest in MAIAC AOD retrievals compared to the output of the MxD04 DT and DB 422 algorithms. Figure 2 shows fire emitted TPM estimates for a series (n=635) of African smoke 423 plumes, as derived from the MxD04 DB AOD product and the MAIAC AOD product, all calculated 424 via multiplication of the plume-integrated AOD anomaly (accounting for pixel area) by the smoke 425 mass extinction coefficient (following Ichoku and Kaufman (2005)). Whilst the MAIAC-derived 426 TPM estimates appear consistent across all VZA's, those from MxD04 BD increase significantly 427 at VZA >  $40^{\circ}$ . Inclusion of plumes observed at high VZA values in the MxD04 DB product used 428 by Mota and Wooster (2018) could lead to artificially high in-plume TPM estimates, and therefore 429 a high bias in the derived C<sub>e</sub> or C<sub>biome</sub> values in FREMv1. When a comparison is made between 430 MxD04\_BD and MAIAC estimated TPM from plumes with VZA < 20° (Fig 2b.), MAIAC-based 431 TPM estimates are typically higher (on average by  $\sim 26\%$ ) than those of MxD04\_BD. This likely

results from the less conservative MAIAC cloud mask and the increased number of valid AODretrievals available over the thickest smoke when compared to MxD04\_BD.





436 (1.5 columns) **Figure 2.** (a) Estimated fire emitted Total Particulate Matter (TPM) in 635 individual 437 smoke plumes, as derived from the 10 km MxD04\_DB AOD product (*orange*) and the 1 km MAIAC 438 AOD product (*blue*), shown as a function of sensor view zenith angle (VZA). (b) Direct comparison 439 of the matching MxD04\_DB and MAIAC-derived TPM values for each plume, restricted to plumes 440 observed at VZA  $\leq$  20°. TPM is calculated from AOD using the equations presented in Ichoku and 441 Kaufman (2005) as described in Section 4.4.

Another update performed in FREMv2 compared to FREMv1 is an improved method for calculating the background AOD value of smoke plumes. FREMv2 applies a localised value for background AOD, as opposed to the large-area-average value applied in FREMv1. The minimum AOD pixel within a buffered area of the smoke plume, in most cases, up-wind of the targeted 446 plume is used as the background AOD value. This approach is similar to that adopted by Ichoku 447 and Kaufman (2005) and Ichoku and Ellison (2014), and justified by the argument that (i) 448 background AODs derived from large scale averaging could be biased by reflectance anomalies 449 or aerosol changes far from the plume, for example dust in the averaging area, and (ii) a large-450 area-averaged background will also be insensitive to the immediate local AOD background of the 451 plume (e.g. during periods of high fire activity when atmospheric particulate matter concentrations 452 are likely to locally be high already).

453 Other adjustments in FREMv2 include the consideration of relative humidity in the estimation of 454 fire generated TPM, expansion to the entirety of continental Africa, and the inclusion of more up-455 to-date land cover and % tree cover information to delineate more precisely the fire-relevant 456 biomes, details of which are included in Section 4.

457

### 458 4. SMOKE EMISSION COEFFICENTS DERIVATION

#### 459 4.1 Geographic Area and Biome Classification

460 For derivation of the biome-dependent FREMv2 smoke emission coefficients ( $C_{biome}$ ) and final 461 emissions estimates, we expanded the Southern Hemisphere Africa (SHAf) region of Mota and 462 Wooster (2018) to include Northern Hemisphere Africa (NHAf). In SHAf the dry season is July to 463 September, and in NHAf November to April, periods which also represent the primary fire seasons 464 in these regions. The dense tropical forests close to the equator (e.g. in northern D.R.C and 465 Gabon) are wetter and less susceptible to large-scale fires compared to those dominated by 466 deciduous and herbaceous vegetation further north and south. These woody savannah and 467 shrubland/grassland areas are those that host most of Africa's biomass burning. In addition to 468 expanding FREMv2 to continental Africa, we also deployed an updated landcover map to provide 21

469 more detailed biome classification. The 2015 European Space Agency (ESA) Climate Change 470 Initiative (CCI) Landcover map (Validated by ESA (2017)) is derived from 300 m spatial resolution 471 PROBA-V observations and comprises 36 landcover type, which we aggregate into five distinct 472 biome classes. Following Mota and Wooster (2018), grassland and woodland savanna are 473 classified as separate biomes as suggested by Korontzi et al. (2004) for fire-related GHG emission 474 reporting. Full details of the CCI land cover class assignment for the FREMv2 biomes can be 475 found in Appendix A. Five biomes were defined by the main vegetation types of *closed canopy* 476 forest, woodland savanna/open forest, grassland, shrubland and managed lands. Since the biome 477 classes of FREMv1 were based on the GLOBCOVER 2009 landcover map 478 (http://due.esrin.esa.int/), which differs from the 2015 CCI landcover map in some respects, the 479 spatial distribution of our biome classes also differ, for example in the Kalahari region of southern 480 Africa. To provide further biome discrimination, FREMv2 also includes use of percentage tree 481 cover (above 5 m height), taken from the 30 m Landsat Vegetation Continuous Fields (VCF) 482 product of 2015 (https://landsat.gsfc.nasa.gov/) (Fig. 2a). Since woodland savanna is by far the 483 largest contributing biome to FRE release over Africa (Appendix B), correct assessment of its 484 smoke emissions is critical to overall accuracy. Areas of woodland savanna having higher tree 485 cover, though still dominated by surface fires (van Leeuwen et al., 2014), have the potential to 486 produce smoke plumes more influenced by surface litter and woody debris (Heil et al., 2010). 487 They may also be more affected by canopy interception of surface-emitted FRP (Freeborn et al., 488 2009; Roberts et al., 2018a). Hence, to improve the precision of C<sub>biome</sub> values for woodland 489 savanna fires, we separated this class into low-woodland savanna and high-woodland savanna 490 using a 20% VCF tree cover threshold.

The FREMv2 biome map was re-projected and aggregated to the Meteosat SEVIRI full disk projection, such that each SEVIRI pixel was assigned a sub-pixel fraction of each biome, with the overall pixel class assigned to the majority fraction (Fig. 2b). Locations of the closed canopy forest

and low- and high-woodland savanna biomes (the main classes having vegetation above 5 m
height) broadly match the % tree cover spatial distribution of VCF product (Fig 2a), and also agree
well with tree cover maps derived previously from MODIS data (e.g. Hansen et al., 2002; Sexton
et al., 2013; Kobayashi et al., 2016).

498



500 (2 columns) Figure 3. (a) Mapped percentage tree cover above 5 metres, as determined from the 501 30 m spatial resolution Landsat Vegetation Continuous fields (VCF) product for 2015. (b) FREMv2 502 biome map for Africa derived from the 2015 ESA CCI Landcover map (itself derived from 300 m 503 PROBA-V imagery) and the Landsat VCF product. Biomes were aggregated from the 36 land 504 cover types defined in the original CCI map, with the two woodland savanna biomes separated 505 using (a) (see Appendix A).

#### 507 4.2 FRP and AOD 2015 Datasets and Fire Matchups

508 The geostationary Meteosat SEVIRI FRP-PIXEL product of Wooster et al. (2015) was a primary 509 input for derivation of the *C<sub>biome</sub>* values. The full spatio-temporal resolution (15 min, 3 km at nadir) 510 FRP-PIXEL product covering the NHAf and SHAf fire season of 2015 was acquired from the 511 EUMETSAT Land Surface Analysis Satellite Applications Facility (LSA SAF: 512 http://landsaf.meteo.pt). The AOD product used was the Terra and Agua combined MODIS 513 MAIAC 550 nm 1 km product (Collection 6 MCD19A2; Lyapustin et al., 2018), described in Section 514 3.

515 Fire activity from the FRP-PIXEL product was assigned to MAIAC AOD measurements of smoke 516 plumes in what we refer to here as 'fire matchup selection'. First, a blob detection (Difference of 517 Gaussian; Lindeberg, 1998) procedure commonly used in computer vision was applied to each 518 MAIAC AOD image to identify regions of interest (ROIs) containing potential smoke plumes. 519 These ROIs were filtered to keep only those having active fires (spatially contiguous FRP pixels 520 observed between 00:00 local time and the MODIS overpass used to produce the MAIAC product) 521 in close spatial proximity. To ensure complete sampling of all fire activity contributing to a given 522 smoke plume, the FRP-PIXEL Quality Product detailed in Wooster et al. (2015) was used to filter 523 out fires that were cloud-obscured leading up to the MODIS overpass time. In cases when both 524 Aqua and Terra MAIAC AOD data were available, the Aqua data acquired closer in-time to the 525 peak of the fire diurnal cycle were preferentially used. This subset of candidate fire matchups was 526 subject to a final manual check to remove any erroneously identified or poorly defined plumes.

527 Each remaining ROI containing a smoke plume had the AOD boundary of the plume defined via 528 histogram thresholding of AOD pixel values (Fig. 4a and 4b). The convex hull of the plume feature 529 was used to define the plume edges and all FRP pixels measured within this bound, between 530 00:00 hrs and the MODIS overpass time relevant to the AOD product, were categorised as being

from the fire which contributed smoke to the plume (Fig. 4c). Though in most of cases FRP pixels were not observed until 08:00 local time. The 1 km resolution of the MAIAC data meant analysis of plume RBG imagery was not needed to help define plume features, unlike with the 10 km MxD04 DB data used in FREMv1 (Mota and Wooster, 2018).

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536

537 (2 column) Figure 4. Example region of interest (ROI) containing the fire shown in Figure 1 along 538 with matchup active fire pixels (white triangles) from the Meteosat FRP-PIXEL product detected 539 between midnight and the MODIS overpass time. (a) MAIAC 1km AOD and SEVIRI active fire 540 (AF) pixels; (b) histogram thresholding to discriminate the plume from the surrounding 'ambient' 541 background. (c) SEVIRI AF pixels detected within the convex hull of the plume are considered to 542 come from the same 'fire' that produced the smoke plume. The fire radiative energy (FRE) of the 543 causal fire is then calculated from these observations and used to match to the AOD-derived total 544 particulate matter (TPM) (see Figure 5).

546 Following Mota and Wooster (2018), only fire matchups for which a single FREMv2 biome 547 represented more than 50% of the observed FRP pixels in a fire were retained. In some cases, 548 the MAIAC AOD cloud mask did screen out some pixels within the smoke plume (e.g. Fig. 1d). 549 though the impact was far less than for the 10 km MODIS MxD04 AOD data. To preserve the 550 accuracy of fire emitted TPM estimates, matchup fires were limited to those with plumes having 551 a MAIAC AOD retrieval at more than 95% of their pixels, and radial bias function interpolation was 552 used to interpolate over the relatively few missing AOD values. To assess the impact of this 553 interpolation on the calculated TPM values, we purposely removed additional AOD pixels and re-554 estimated their value via the same process. Minimal impact was shown due to the quality of the 555 interpolation procedure and the fewer than 5% of in-plume AOD pixels it was required to be 556 applied to.

557 After pre-processing and data screening, 968 fire matchups remained for C<sub>biome</sub> derivation, and 558 these are mapped in Figure 6d. Each had its FRE and column integrated mass of total plume 559 particulate matter estimated, the former from the temporal integration of FRP from the start of fire 560 activity on that day to the time the MAIAC AOD data were acquired. The latter was calculated 561 from the plumes' total plume AOD anomaly divided by the smoke aerosol mass extinction 562 coefficient,  $\beta_e$  (in m<sup>2</sup>.g<sup>-1</sup>) and multiplied by the total plume area (m<sup>2</sup>) calculated over all AOD pixels 563 (following Ichoku and Kaufman, 2005).  $\beta_e$  can be measured in situ and its value depends on 564 several factors including relative humidity, age of the smoke, vegetation burned and the 565 wavelength used in measurements. Reid et al. (2005) provide a detailed review of values for  $\beta_{e}$ 566 in smoke plumes that range between 3.8 - 4.5 m<sup>2</sup>.g<sup>-1</sup>, and combining these with the values of 567 Abel et al. (2005) for South African fires that range from 2.22 - 3.37 m<sup>2</sup>.g<sup>-1</sup> we assume an 568 intermediary  $\beta_e$  of 3.5±1.0 m<sup>2</sup>.g<sup>-1</sup>.  $\beta_e$  values for smoke have been shown to increase with aerosol 569 ageing (Formenti et al., 2003; Abel et al., 2005) and also with relative humidity (RH) (Chin et al., 570 2002; Koppmann et al., 2005). The RH for each of our plumes was taken from the ERA-Interim 26

571 reanalysis at 762 m altitude (Balsamo et al., 2015), and all matchup plumes showed coincident 572 RH values < 70% so we assume minimal effect of RH on  $\beta_e$  since below this RH threshold inflation 573 of  $\beta_e$  for smoke generated from biomass burning is typically less than 10% (Chin et al., 2002; Reid 574 et al., 2005). The  $\beta_e$  of fresh smoke aerosol has been shown to increase by 20 to 50% as it ages 575 after 1 - 4 days (Reid et al., 1998; Abel et al., 2005). However, for the majority of our fire matchups, 576 significant fire activity was not observed until 08:00 hrs or later in the day, resulting in the oldest 577 smoke in our plumes being around 7 or 8 hours old and thus limiting the extent to which  $\beta_e$  grows 578 due to plume ageing. We therefore retain the  $\beta_e$  value of 3.5±1.0 m<sup>2</sup>.g<sup>-1</sup> used by Mota and Wooster 579 (2018), and the uncertainty range attached to this value also includes these potentially higher 580 bounds. We do however recommend further investigation into the potential for biome-dependent 581 variability in  $\beta_e$ .

## 582 4.3 Derivation of Smoke Emission Coefficients (C<sub>biome</sub>)

583 Our fire matchups were used to derive a set of smoke emission coefficients (Table 1; Fig. 5) for 584 each of the biomes defined in Section 4.1, based on zero-intercept linear orthogonal distance 585 regression (ODR). Uncertainties in each variable are accounted for in ODR and are calculated 586 from the combined AOD uncertainty measures provided in the MAIAC product, the uncertainty in 587  $\beta_e$ , and the FRP uncertainties provided in the FRP-PIXEL product (Wooster et al., 2015).



(2

590 *column*) **Figure 5.** Smoke emission coefficients ( $C_{biome}$ ; in g.MJ<sup>-1</sup>) for the six African fire-affected biomes defined in Section 4.1, each derived from the slope of an orthogonal distance regression 591 592 (ODR) between data on fire-emitted total particulate matter (TPM) and matching fire radiative 593 energy (FRE). Grey shaded area defines the 95% probability prediction interval of the ODR-594 derived slope. Each scatterplot is accompanied by an illustrative insert that depicts the typical 595 landcover for the biome as seen in Google Earth (example locations are Closed Canopy Forest, 596 10.359° S, 19.086° E; Grassland 21.180° S, 19.560° E; Managed Land 10.495° N, 7.586° E; Low-597 Woodland Savanna 7.085° N, 27.095° E, High-Woodland Savanna 12.523° S, 23.323° E; 598 Shrubland 23.055° N, 22.242° E).

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**Table 1.** Biome-dependent smoke emission coefficients ( $C_{biome}$ ) and accompanying uncertainties for the African fire-affected biomes mapped in Figure 3b and calculated from FREMv2 and from FREMv1 (reported in Mota and Wooster (2018)). Matching FEER-equivalent coefficients are also shown, based on the geographical location of fire matchups and the FEER Ce (1° × 1°) product of Ichoku and Ellison (2014).

Biome	FREMv2 C <sub>blome</sub> (g MJ <sup>-1</sup> )	FREMv1 C <sub>biome</sub> (g MJ <sup>-1</sup> )	FEER equivalent C <sub>biome</sub> (g MJ <sup>-1</sup> )
Closed canopy Forest	34.33±2.85	65.63±0.91	16.34
Managed land	13.98±1.03	15.62±0.34	15.80
Grassland	9.99±0.29	13.03±0.23	10.98
Shrubland	12.17±0.42	17.36±1.06	10.97
Low-woodland savanna	12.10±0.30	19.75±0.49	12.78
High-woodland savanna	16.43±0.32	19.75±0.49	13.81

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608 Our FREMv2 smoke emission coefficients were compared to those of FREMv1 (Mota and 609 Wooster, 2018) and FEER (Ichoku and Ellison, 2014; Table 1). In FEER,  $C_e$  values are derived 610 for each individual 1° grid cell using the MODIS archive of that cell, rather than by biome. To 611 derive FEER-equivalent  $C_{biome}$  values for comparison, we intersected all FREMv2 fire matchups 612 with each FEER 1° grid cell and then calculated the mean FEER  $C_e$  for each of our biomes.

Our FREMv2  $C_{biome}$  values range from 9.99 ± 0.29 g MJ<sup>-1</sup> (grassland) to 34.33 ± 2.85 g MJ<sup>-1</sup> 613 614 (closed canopy forest), and in all biomes are lower than those of FREMv1 with an average 615 decrease of 27% across all biomes, that for closed canopy forest is almost halved compared to 616 FREMv1. These differences relate to a combination of the far more appropriate MAIAC AOD 617 product, the improved method for background AOD estimation, the updated and extended biome 618 mapping, and the inclusion of fire matchups from NHAf as well as SHAf. Since plume TPM values 619 are most strongly influenced by the AOD product used, use of MAIAC AODs is likely to have the 620 most significant impact on the C<sub>biome</sub> values and the product characteristics can both increase 621 and decrease these compared to use of MxD04. As discussed in Sections 2.3 and Section 3, 30

622 MxD04 AOD retrievals are VZA dependent and bias can be seen at the swath edges (Sayer et 623 al., 2015), meaning TPM estimates for plumes observed at wide VZA can be significantly inflated 624 (Fig. 2a). In the original FREMv1 methodology such plumes were retained during the C<sub>biome</sub> 625 derivation, and likely resulted in overestimated TPM values and thus FREMv1 C<sub>biome</sub> values. 626 Additionally, the MAIAC AOD product uses an area-weighted method to grid L1B MODIS pixels 627 to a 1 km pixel size prior to AOD retrieval, providing finer detail and improved plume distinction at 628 VZA > 20°, allowing inclusion of smaller fires in the matchup dataset. A feature of the MAIAC 629 AOD product likely to increase estimates of TPM and thus FREMv2 C<sub>biome</sub> values is its less 630 conservative cloud mask, which resulted in MAIAC-derived TPM estimates being ~ 26% higher 631 than the equivalent MxD04\_BD values when only plumes observed at low VZA's are considered 632 (Fig. 2b).

633 Final Chiome values for shrubland and low-woodland savanna are close (within 0.07 g MJ<sup>-1</sup>) (Table 1), indicating these biomes appear broadly equivalent with respect to their fire and particulate 634 635 matter emissions characteristics. The number of fire matchups identified in both the closed 636 canopy forest and managed land biomes are significantly fewer than for the other biomes, largely 637 because of fewer identifiable fires and the fact that many do not meet the matchup criteria. The 638 lower number of matchups mean the C<sub>biome</sub> coefficients for these two biomes have significantly 639 larger uncertainties ( $3 \times$  and  $10 \times$  higher than for the other biomes respectively), and their smoke 640 emissions coefficients are quite strongly influenced by a relatively few high FRE fire matchups. 641 As discussed in Section 2.1, the SEVIRI FRP-PIXEL product has a minimum AF detection limit of 642 around 30 - 40 MW, and fires burning below this will often remain undetected until they breach 643 this threshold or other fires within the same pixel cause the pixels FRP total to exceed this. 644 Agricultural fires in managed lands are typically guite small (Zhang et al., 2017), and in closed 645 canopy forests surface fires burning on the forest floor will be partially obscured by the tree canopy 646 (Roberts et al., 2018a). Both these effects will result in a lower FRE total being measured at a fire

647 than might otherwise be the case, and these biomes also have the highest percentage of 648 matchups with FRE below 1×10<sup>7</sup> MJ. But to the extent that these issues affect both the matchup 649 fires of these biomes and the overall set of SEVIRI-detected fires to which the coefficients are 650 applied, the effect of these issues is taken into account. FREMv2 coefficients are on average only 651 18% higher than FEER-equivalent  $C_{biome}$  values, with FREMv2  $C_{low-woodland savanna}$  (which has 652 one of the largest matchup sample sizes) showing very good agreement with its FEER- equivalent (12.10±0.32 and 12.78 g MJ<sup>-1</sup> respectively). FREMv2  $C_{low-woodland \ savanna}$ ,  $C_{shubland}$  and 653 Cgrassland values are also all within 11% of their FEER-equivalents. Cclosed canopy forest shows the 654 655 greatest divergence at around twice the FEER-equivalent, perhaps stemming from the relatively 656 small sample size (n=38) and/or the high variability in FEER  $C_e$  values for this biome (Fig. 6b).

657 The newly derived FREMv2 C<sub>biome</sub> values of Table 1 were used to generate a SEVIRI per-pixel 658 smoke emissions coefficient ( $C_e$ ) product for Africa, for subsequent use in smoke emissions estimation. Ce values for each SEVIRI pixel across the African continent were calculated based 659 660 on the weighted mean of the relevant C<sub>biome</sub> values and the per-pixel biome fractional cover 661 derived in Section 4.1. Figure 6a shows the resulting FREMv2 C<sub>e</sub> product, along with the FEER 662  $C_e$  (1° × 1°) product (Ichoku and Ellison, 2014) (Fig 6b) and their difference (averaged to 1° 663 resolution) (Fig. 6c). As previously described, FEER  $C_e$  values are calculated for each 1° grid cell 664 from matchup fires within that cell, rather than per biome, so the spatial variability of FEER  $C_e$ 665 values is far higher than for the biome-driven FREMv2 Ce values of Fig. 6a. Across Africa, FEER 666  $C_e$  are on average higher than those of FREMv2 by 0.53 g.MJ<sup>-1</sup>, and this difference is dominated 667 by regions where the FEER coefficient is significantly higher than that of FREMv2 (Fig. 6c). The 668 most similar regions are generally those well sampled by fire matchups in FREMv2 (Fig. 6d), 669 whereas those with the highest differences have few or no fire matchups in FREMv2, due primarily 670 to relatively low fire activity being recorded there but also frequent cloud cover in 2015. In some 671 cases such areas appear to exhibit notably high FEER  $C_e$  values (> 40 g.MJ<sup>-1</sup>). Future work on 32

672 FREM will extend the fire matchup sampling period to multiple years to obtain additional matchups 673 in regions currently relatively poorly sampled, and will aim to better understand the large divergence in  $C_e$  values compared to FEER in these locations. 674



677 (2 columns) Figure 6. (a) FREMv2 smoke emission coefficient ( $C_e$ ) mapped at 0.05°; (b) the 678 matching 1° FEER C<sub>e</sub> product; and (c) the difference. In (c), FEER grid cells whose value was 679 derived from gap-filling or which were calculated from less than 15 samples were removed (see 680 Ichoku and Ellison (2014) for full details). (d) Shows the spatial distribution of the fire matchups 681 used to derive the FREMv2 C<sub>biome</sub> values of Figure 5.

### 682 5. FREMV2 EMISSION INVENTORY DEVELOPMENT

#### 683 5.1 Total Particulate Matter (TPM) Emissions

684 We used the FREMv2 Ce product of Figure 6a to convert the 2013 to 2018 Meteosat FRP-PIXEL 685 product (Wooster et al., 2015) into the highest spatio-temporal resolution TPM emissions inventory yet available over Africa (15 min, 3 km at the sub-satellite point). We compared these 686 687 emissions to a series of other inventories widely used by the research and operational 688 communities (Figs. 9 and 10). Hourly averages of SEVIRI FRP (in MW) multiplied directly by the 689 FREMv2  $C_e$  product provide mapped instantaneous TPM emission rates (in kg.s<sup>-1</sup>), whereas 690 multiplication by FRE gives emission totals for the defined FRP temporal integration period. 691 Figure 7 presents TPM emissions estimates for FREMv2, GFASv1.2 (Kaiser et al., 2012; 692 https://apps.ecmwf.int/datasets/data/cams-gfas/), GFEDv4.1s (Van der Werf et al., 2017; 693 www.globalfiredata.org/), along with FEERv1.0-GFAS1.2 (the FEER  $C_e$  product applied to 694 GFASv1.2 FRP estimates, Ichoku and Ellison, 2014; www.feer.gsfc.nasa.gov/data/emissions/) 695 and FEERv1.0-SEVIRI (the FEER  $C_e$  product applied to SEVIRI FRE).

696 FREMv2 estimates mean annual TPM emissions for Africa at 32.41±1.86 Tq.yr<sup>-1</sup> for the five years 697 studied, 38% of which are generated from fires in NHAf, and 62% SHAf. Pan-African totals are 698 114% and 69% higher than GFASv1.2 and GFEDv4.1s respectively, with estimates from SHAf 699 fires showing greater divergence in both cases. The FREMv2 values however, are within 11% of 700 the top-down inventories FEER-GFASv1.2 and FEER-SEVIRI, agreeing with Ichoku and Ellison 701 (2014) who also show FEER-GFAS to be higher than GFASv1 and GFEDv3 by similar factors 702 over NHAf and SHAf. Kaiser et al. (2012) report that GFASv1.2 smoke aerosol emissions must 703 be multiplied by a global scaling factor of 3.4 before input into atmospheric models if they are to 704 provide modelled AODs more in line with observations. This provides evidence that Africa's TPM 705 emissions are indeed higher than GFAS (and GFED) currently estimate. As demonstrated by 34

706 Wooster et al. (2015) and Hall et al. (2019), at the time of their overpasses, polar-orbiter based 707 AF products detect more "small" (i.e. low FRP) fires compared to the SEVIRI FRP-PIXEL product 708 as a result of their finer pixel size and thus lower minimum FRP detection limit. This effect is 709 amplified at SEVIRI VZA > 40° due to the growth of sensors pixel footprint area. However, it was 710 also shown that over the course of several days, the far more frequent data available from SEVIRI 711 allows an increased number of AF detections overall (Wooster et al., 2015; Hall et al., 2019). 712 Since the SEVIRI-derived FREMv2 emissions agree well with those of FEER, which are derived 713 from MODIS FRP data (which have a significantly lower minimum FRP detection limit than 714 SEVIRI; Roberts et al., 2015), our findings suggest that FREMv2 accounts for the smoke emission 715 contribution from a substantial proportion of the active fire pixels remaining undetected by SEVIRI. 716 This is likely because, whilst SEVIRI fails to detect these lowest FRP fires, the smoke they 717 generate has contributed to the AOD in the 1km MAIAC product used to generate the biome-718 dependent smoke emissions coefficients from the fire matchups. Thus, our FREMv2 Chiome and 719  $C_e$  data contain an inherent 'boost' from the TPM emitted from undetected low-FRP active fire 720 pixels. The 1 km spatial resolution of the MAIAC AOD, which enables us to distinguish and use 721 many more of the smaller smoke plumes than the higher resolution MxD04 product, helps enable 722 this compared to the 10 km AOD data used in FREMv1.


(2 columns) Figure 7. Monthly total particulate matter (TPM) emissions from landscape fires for
2013 to 2018, as derived using the FREMv2 methodology *(blue)* applied to the Meteosat FRPPIXEL product of Wooster et al. (2015). Corresponding monthly TPM emissions are shown from
GFEDv4.1s *(green)*, GFASv1.2 *(purple)*, and the FEERv1.0 coefficients applied to the GFASv1.2 *(red)* and SEVIRI FRE data *(yellow)*.

730 The five inventories of Figure 7 show similar temporal patterns of TPM emission, with clearly 731 identifiable NHAf and SHAf burning seasons and annual peaks and minima occurring in the same 732 years. Like Mota and Wooster (2018), FREMv2 predicts an earlier peak in TPM emissions over 733 SHAf than do the other inventories. Peak smoke emission in FREMv2 occurs in July in every year 734 but one, whereas in all other inventories (except FEER-SEVIRI, which has peak TPM split equally 735 between July and August) this occurs in August. A potential cause may be the splitting of high 736 and low woodland savanna biomes in FREMv2. High-woodland savanna is mainly concentrated 737 just below the equator (Fig. 3b), in a region that burns far earlier than the dominant southern 738 African fire season. Therefore, due to FREMv2 discriminating between high and low woodland 739 savanna and grasslands, the dominant  $C_e$  values applied to the SEVIRI FRP-PIXEL data vary not 740 only by biome, but also over time due to the seasonal progression of fire activity across the 741 continent. Due to the FREMv2 biome map being based on land cover and VCF data from 2015, 742 changes in land cover in the 2013 to 2018 period are not accounted for. However, canopy 743 changing crown fires are very rare in low and high-woodland savanna, shrublands and grasslands 744 - which are dominated by surface fires (Van Wilgen et al., 1990; Heil et al., 2010; Van Leeuwen 745 et al., 2014) - and so fire-related changes in tree canopy cover is very limited in these biomes 746 (Zhou et al., 2019). However, year-to-year anthropogenically driven landcover changes could 747 impact the methodology and it may become more important to account for such changes within 748 our FREMv2 coefficients as this increases. Future FREM versions will include more regular 749 landcover updates in the mapped emissions coefficient estimation.

The similar TPM estimates generated by FEER-GFASv1.2 (which uses MODIS FRP) and FEER-SEVIRI (and FREMv2, which both use SEVIRI FRP), compared to the far lower values of GFASv1.2 (also based on MODIS FRP) indicate that the higher TPM emissions of the top-down approaches stem dominantly from the  $C_e$  values applied to FRP measures, and less so from the

754 source of the FRP observations used. That said, in the case of FEER-GFASv1.2 it is noteworthy 755 that whilst the FEER C<sub>e</sub> values are derived from use of direct MODIS FRP observations, the 756 MODIS FRP data used within GFAS undergoes several stages of processing and thus presents 757 different FRP values to those originally provided by the MODIS MOD14/MYD14 products (Kaiser 758 et al., 2012). The similarity of the TPM emissions estimates generated by FREMv2 and the two 759 FEER inventories supports the case that higher emissions estimates come mainly from the 760 magnitude of Ce values, since the FREMv2 and FEER Ce values agree well in general (Table 1 761 and Fig. 8). Future work will evaluate the quality of the final TPM emissions generated by FREMv2 762 by placing them as inputs to atmospheric chemical transport models and comparing the resulting 763 aerosol fields to independent data such as AERONET and surface PM<sub>2.5</sub> measures.

764 We generate 2016 FREMv2 TPM emissions at 0.05° across the African continent, a spatial 765 resolution x5, x2 and x2 times higher than offered by GFEDv4.1s, GFASv1.2 and FEER-766 GFASv1.2 respectively. We present these for Africa along with a magnified 20×20° region in 767 Figure 8. The FEER  $C_e$  product with its 1° grid cell resolution can, in theory, be applied to the 768 native SEVIRI FRP-PIXEL product as has been done for the FEER-SEVIRI dataset shown in 769 Figure 8. However, the spatial resolution of the FEER  $C_e$  product is 20 times lower than that of 770 the FRP observations, so these FEER-SEVIRI derived emissions estimates do not account for 771 the finer detail inter-biome spatial variations that FREMv2 does. Additionally, the appropriateness 772 of applying FEER C<sub>e</sub> values, which are derived from 'raw' MODIS FRP data, to 'raw' SEVIRI FRP 773 data is unclear, particularly when their quite different minimum-FRP detection limits are 774 considered.

In general, the spatial distribution of the African TPM emissions is somewhat similar across all
five inventories. In line with the temporal trends of Fig. 7, we see notably higher TPM totals for
the three top-down emission inventories, particularly FREMv2. The high spatial resolution of 38

FREMv2 provides more precisely detailed spatial information on smoke emissions than do the other inventories, which maybe relevant for supporting improved local scale air quality modelling. In FREMv2, deriving separate  $C_{biome}$  coefficients for high- and low-woodland savanna results in distinctly higher total TPM emissions in north Angola/south D.R.C. compared to the other inventories. This region is dominated by high-woodland savanna (Fig. 3b), and demonstrates the impact of the more spatially resolved biome classification used in FREMv2.



785 (1.5 columns) **Figure 8.** Total particulate matter (TPM) emission density (g.m<sup>-2</sup>) across Africa for 786 2016 as determined by GFEDv4.1s (0.25° grid cells), GFASv1.2 (0.1° grid cells), FEERv1.0-787 GFASv1.2 (0.1° grid cells), FEERv1.0-SEVIRI (0.05° grid cells), and FREMv2 inventory derived 788 herein (0.05° grid cells). The red 20°  $\times$  20° region outlined in the top left GFED plot is shown 789 magnified for each inventory in the right-hand column.

790 The distribution of biomes across NHAf and SHAf has a significant impact on the contribution 791 each biome has to the total TPM emissions of each hemisphere (Fig. 9). Both closed canopy 792 forest and grassland show a similar percentage contribution to total TPM emissions in each 793 hemisphere, with mean annual emissions totals within 2% of each other and a combined mean 794 contribution of 17% and 20% of total TPM emissions for NHAf and SHAf respectively. In NHAf, 795 fires from managed lands and shrublands contribute most to annual TPM emissions, though their 796 fractional contributions exhibit significant seasonal variations (45% and 38% across the year 797 respectively). In SHAf these two biomes show a narrower range across the year and an overall 798 lower contribution to total TPM emissions, which are instead dominated by high-woodland 799 savanna fires between May and November. In both NHAf and SHAf, the highest monthly 800 contribution of emissions from managed lands occurs outside the primary burning season (Nov-801 April in NHAf and July-Sept in SHAf), potentially due to deliberate post-harvest or end-of-growing 802 season burning (Yevich and Logan, 2003).

803



*(1.5 columns)* Figure 9. Mean monthly contribution (%) of fires to total particulate matter (TPM)
emissions in each of the six FREMv2 biomes from 2013 to 2018 in (a) Northern Hemisphere Africa
and (b) Southern Hemisphere Africa.

808 5.2 Trace gas and total carbon emissions

As with FREMv1 (Mota and Wooster, 2018), trace gas emissions estimates are derived from
FREMv2 outputs via application of the standard gaseous emission factors (EFs) of Andreae and
Merlet (2001), which are also used in GFEDv4.1s, GFASv1.2 and FREMv1. Although EF updates
to Andreae and Merlet (2001) have recently become available (Andreae, 2019), we maintain the
42

813 use of the original EFs for consistency and ease of comparison with the other fire emissions 814 databases. Unlike the bottom-up approaches in which the EFs are applied to total fuel 815 consumption estimates, in the FREM methodology trace gas emissions are estimated directly 816 from the observed FRP or FRE values using a set of trace gas emissions coefficients. The 817 approach is similar to that of Huijnen et al. (2016) who estimated CO<sub>2</sub> and CH<sub>4</sub> emissions from 818 fire-emitted CO estimates. The trace gas emission coefficients in FREM are calculated using the 819 EF ratios between the relevant gas and TPM, which are then multiplied by the TPM C<sub>biome</sub> values 820 presented in Section 4.4 (Table 1) (Mota and Wooster, 2018):

821 
$$C_{gas}^{biome} \left[ g. MJ^{-1} \right] = \frac{EF_{gas}^{biome} \left[ g. kg^{-1} \right]}{EF_{TPM}^{biome} \left[ g. kg^{-1} \right]} \cdot C_{TPM}^{biome} \left[ g. MJ^{-1} \right]$$

822 These coefficients are applied directly to the geostationary FRP or FRE data to estimate the trace 823 gas emissions. Trace gas emissions coefficient derivation for each SEVIRI pixel used an area-824 weighted mean of the biome-specific EF ratios (similarly to  $C_e$  product derivation in Section 4.3), 825 thereby generating an emission coefficient map for each gas. The biomes used by Andreae and 826 Merlet (2001) to report their EF values are less detailed than those of FREMv2, resulting in all but 827 the closed canopy forest biome using the same EF values. This relative lack of EF detail affects 828 all the fire emission inventories compared herein, and points to a potential need for more research 829 focused on more finely detailed EFs with respect to vegetation type.

Final monthly CO<sub>2</sub>, CH<sub>4</sub> and CO emissions are shown in Figure 10, which exhibit a similar seasonal pattern to those of TPM (Fig. 7), and mean annual totals are shown in Table 2. Direct retrieval of CO atmospheric concentrations is carried out using data from instruments such as MOPITT (Worden et al., 2010) and TROPOMI (Veefkind et al., 2012), and comparisons of GFEDv2 and GFEDv3 CO emissions with MOPITT-derived CO have previously suggested that GFED underestimates fire emitted CO over Africa by up to 50% (Chevallier et al., 2009; Kopacz 836 et al., 2010; Pechony et al., 2013). Comparisons of the different GFED versions shows that for 837 NHAf and SHAf, GFEDv4.1s CO emissions were around 30% lower than GFEDv3 (Van Der Werf 838 et al., 2017), pointing to significant underestimation of CO emissions by GFEDv4.1s over Africa, 839 similarly to that of TPM. Studies showing African burned area to be far higher when mapped using 840 20 m Sentinel-2 MSI imagery than with the 500 m MCD64A1 product used in GFED supports this 841 idea (Tsela et al., 2014; Hawbaker et al., 2017; Roteta et al., 2019). The substantially higher CO 842 emissions provided by FREMv2 and by the other top down approaches compared herein may 843 therefore be more realistic than the lower values provided by the bottom up inventories.



*(2 columns)* Figure 10. Monthly total emissions (Tg) of (a) CO<sub>2</sub>; (b) CO; and (c) CH<sub>4</sub> for African
landscape fires as estimated by FREMv2 (*blue*) between 2013 and 2018. Corresponding values
from GFEDv4.1s (*green*), GFASv1.2 (*purple*), FEERv1.0-GFASv1.2 (*red*) and FEERv1.0-SEVIRI
(*yellow*) are shown for comparison.

850 FREMv2 emissions of total carbon were calculated from the summed carbon contents of the CO<sub>2</sub>, 851 CO and CH<sub>4</sub> emissions, which typically contribute more than 99% of total carbon release in 852 savanna/grassland and tropical forest fires (Andreae, 2019). Estimates of total fuel consumption 853 in terms of dry matter consumed (DMC) were then calculated on the assumption of a 50% DM 854 fuel carbon content (Van Der Werf et al., 2010; 2017), with mean annual totals reported in Table 2. 855 Table 2. Mean annual total carbon and trace gas emissions for 2013 to 2018, along with dry 856 matter consumed (DMC) totals (Tg.yr<sup>-1</sup>), for Northern and Southern Hemisphere Africa as 857 estimated by the different fire emissions inventories compared herein including FREMv2 (final 858 column).

859

	Mean annual emissions (Tg yr-1)										
	GFAS		GFED		FEER-GFAS		FEER-SEVIRI		FREMv2		
Species	NHAF	SHAF	NHAf	SHAf	NHAf	SHAf	NHAf	SHAf	NHAf	SHAf	
с	346.3	470.3	398.1	651.4	588.3	1007	593.5	1023	682.7±76.9	1090±46	
CO2	1188	1619	1368	2242	2018	3457	2045	3525	2349±270	3751±167	
со	48.25	62.87	54.74	87.62	82.54	141.03	77.94	134.8	91.52+10.53	147.1±6.6	
CH4	2.033	2.421	1.921	2.953	3.375	5.708	2.966	5.164	3.623±0.417	5.894±0.263	
Dry Matter Consumed	692.6	940.6	796.2	1302.8	1176.6	2014	1187	2046	1365±154	2180±92	

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## 862 5.3 Fuel Consumption per unit area $(F_c)$ from 20 m Burnt Area

As discussed in Section 2.1, poorly constrained fuel load and combustion completeness parameters contribute a significant part of the uncertainty in bottom up estimates of fuel consumption per unit area ( $F_c$ , in kg.m<sup>-2</sup>), and unmapped small burns missed by the 500 m 866 MCD64A1 MODIS burned area (BA) product adds further uncertainty and possibly bias into the final derived fuel consumptions (Reid et al., 2009; Vermote et al., 2009; Van Leeuwen and Van 867 868 Der Werf, 2011; Tsela et al., 2014; Hawbaker et al., 2017; Roteta et al., 2019). As first 869 demonstrated by Mota and Wooster (2018), it is however possible to generate FREM-derived 870 estimates of F<sub>c</sub> through an inversion of the original Seiler and Crutzen (1980) BA-based approach 871 to fire carbon emissions estimation. The 500 m MODIS MCD64A1 BA product used in GFED (Van 872 Der Werf et al., 2017) and in the calibration of GFAS (Kaiser et al., 2012) was used in mapping 873 of F<sub>c</sub> across southern Africa in FREMv1 (Mota and Wooster, 2018), but here we use the far higher 874 spatial resolution (20 m) FireCCISFD11 African BA product for 2016 that has been shown to map 875 up to 60% more BA in some areas through the detection of far smaller burn patches (Roteta et 876 al., 2019). We mapped  $F_c$  across Africa at 0.25° (Fig. 11a) by gridding the FREMv2 dry matter 877 consumed (DMC) data for 2016 (Section 5.2) and dividing this by the matching BA calculated 878 from the approximately 1.5×10<sup>6</sup> potential FireCCISFD11 20 m pixels falling in each 0.25° grid cell. 879 Note that these per-pixel  $F_c$  values apply only to the burned area patch inside a given pixel, and 880 not the 0.25° pixel as a whole. Compared to  $F_c$  provided by the model-based GFEDv4.1s (Figure 881 11b), around half the grid cells show a significantly higher  $F_c$  value in FREMv2, which is to be 882 expected due to the overall higher carbon emissions of the former (Fig. 10 and Table 2), and for 883 the remaining cells the two inventories provide similar values (see difference map in Appendix C). 884 There are some unprocessed tiles in the FireCCISFD11 product (see Roteta et al. (2019)), 885 resulting in a small minority of  $0.25^{\circ}$  grid cells having unreported or unrealistically high  $F_c$  values 886 in excess of 10 kg.m<sup>-2</sup>. Avoiding use of these few anomalous cells, we derived biome-specific 887  $F_c$  frequency distributions and statistics based on those cells where the biome covered at least 888 80% of the cell area and where more than 5% of the cell burned (Figure 11c). Distributions are 889 heavy-tailed and show a spatial variability somewhat similar to that derived by Roberts et al. 890 (2011) using an alternative FRE and BA based approach. There are relatively few fuel

consumption databases derived from field measurements, but van Leeuwen et al. (2014) provides
summary statistics for African savannah burns and our low woodland savannah and grassland
averages listed in Figure 11c are very close to their 0.34 kg.m<sup>-2</sup> mean.



896 (2 columns) **Figure 11.** Fuel consumption per unit area ( $F_c$ , kg.m<sup>-2</sup>) mapped at 0.25° from (a) 2016 897 FREMv2 dry matter consumed (DMC) totals and the FireCCISFD11-estimated burned area, and 898 (b) GFEDv4.1s. (c) Per-biome FREMv2  $F_c$  frequency distributions and derived means, medians 899 and standard deviations. Note that  $F_c$  values in (a) apply only to the burned area patch inside a 900 given pixel, and not the 0.25° pixel as a whole.

901 *6.* 

## SUMMARY AND CONCLUSIONS

We have provided significant advances to the fully top-down 'Fire Radiative Energy Emissions' (FREM) landscape fire emissions methodology of Mota and Wooster (2018), and have used this to develop the highest spatio-temporal resolution African landscape fire emissions inventory currently available. The approach will form the basis of a new fire emissions product to be delivered by the EUMETSAT Land Surface Analysis Satellite Application Facility (http://landsaf.meteo.pt), and will in future be extended back to 2004 using the full Meteosat SEVIRI FRP archive already exploited to study African fires by Roberts et al. (2018b).

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910 The FREM approach bypasses the total fuel consumption step of 'bottom up' fire emissions 911 methodologies, recognised as a key source of uncertainty (Reid et al., 2009; Vermote et al., 2009; 912 Van Leeuwen and Van Der Werf, 2011). The method generates smoke emissions estimates 913 directly from satellite retrievals of FRP and relies on a set of biome-dependent smoke emission 914 coefficients ( $C_{biome}$ , g.MJ<sup>-1</sup>) that are here derived from almost one thousand matchups between 915 Meteosat SEVIRI Fire Radiative Energy (FRE) estimates (Wooster et al., 2015) and in-plume total 916 particulate matter (TPM) estimates made using the 1 km MAIAC AOD product (Lyapustin et al., 917 2018). Our FREMv2 methodology introduces significant improvements to all stages of the original 918 FREM approach, and most particularly those associated with the prior use of the 10 km MODIS 919 MxD04 AOD product for TPM estimation. Issues with use of MxD04 include, masking of thick smoke as cloud (Levy et al., 2013; Raffuse et al., 2013; Remer et al., 2013; Livingston et al., 920 921 2014), the impact of sensor VZA dependence on AOD retrieval (Saver et al., 2015; Mhawish et 922 al., 2019), and the products' relatively low spatial resolution making plumes harder to discriminate 923 and requiring a focus on the largest fires when deriving the FREM coefficients (Lyapustin et al., 924 2011; Raffuse et al., 2013; Remer et al., 2013; He et al., 2017). Our new use of the 1 km MAIAC 925 AOD product of Lyapustin et al. (2018) addresses each of these issues, and we also derive a 926 more up to date and detailed fire-relevant mapping of pan-African biomes using the CCI Land 927 Cover 2015 (ESA, 2017) and Landsat-derived percentage tree cover information. The latter 928 enables improved specification of the smoke emission coefficients in the woodland savanna 929 biome, which annually contributes the most to Africa's fire radiative energy release. Expansion 930 of the FREMv2 inventory to include both NHAf and SHAf enabled many more FRE-AOD matchup 931 fires to be included in the smoke emissions coefficient generation compared to FREMv1, including 932 many more small fires whose plumes can be discriminated using the 1 km MAIAC AODs. This 933 also enabled localised values of plume background AOD to be selected in a more representative 934 manner than in FREMv1 (Mota and Wooster, 2018), and the impact of relative humidity on our 935 use of smoke mass extinction coefficient values was also assessed during our methodology.

936

The evolutions reported herein result in a set of FREMv2 biome-dependent smoke emission coefficients (*C*<sub>biome</sub>) for total particulate matter 10% to 48% lower than those of FREMv1 (Mota and Wooster, 2018). The FREMv2 coefficients are now closer to those of another top-down FRPbased emissions methodology (FEER; Ichoku and Ellison 2014) compared to those of FREMv1, particularly for the low-woodland savanna, shrubland and grassland biomes which are each considerable sources of smoke emissions (Table 1). Some significant differences between the coefficients in FREMv2 and FEER do remain, but mostly in regions showing relatively few fires

944 and thus which are rather poorly sampled by FREMv2 fire matchups. A significant advantage of 945 FREMv2 over FEER is that the spatial mapping of the final  $C_e$  smoke emission coefficients across 946 Africa is easily derived from the five FREMv2 C<sub>biome</sub> values reported in Table 1 and the SEVIRI-947 pixel area biome coverage, meaning that FREMv2  $C_e$  updates related to landcover changes can 948 be easily calculated using only an updated landcover map. In the FEER methodology of Ichoku 949 and Ellison (2014), C<sub>e</sub> are mapped on a 1° grid cell basis from fire matchups observed within each 950 cell using more than a decade of data from the MODIS record, and landcover change related 951 updates thus require complete collection of a set of new fire matchups and re-derivation of  $C_e$ .

952

Whilst our FREMv2 approach addresses the principal uncertainties and biases in the original
FREMv1 (and indeed the FEER) methodologies, there remain sources of uncertainty and
limitations that will benefit from further investigation. These include:

956 i) Elucidation of the effect of small fires having an FRP below the minimum geostationary 957 active fire detection limit, and the extent to which these are now accounted for via the 958 ability to include smaller fire matchups during the  $C_{biome}$  derivation via use of the 1km 959 MAIAC AOD product (Section 5.1)

960 ii) Further investigation of smoke mass extinction coefficients ( $\beta_e$ ) used to estimated column-961 integrated TPM from AOD, and the relevance of smoke ageing over the period relevant 962 here (typically < 8 hrs).

963 iii) Assessment of the impact of the exponential growth of the geostationary sensor pixel area
964 far from nadir, which amplifies the non-detection of low FRP fires by raising the minimum
965 FRP detection limit (Wooster et al., 2015).

966 iv) Improvement of the smoke emissions coefficients for closed canopy and managed land,
967 which are currently derived from a relatively small number of matchup fires compared to
968 the other biome classes.

970 With regards to point (iii), in the scope of any future pan-tropical and mid-latitude fire emissions 971 product this issue can be partly overcome by applying the FREMv2 methodology to a suite of 972 geostationary satellite FRP data, to derive a global  $C_e$  product which could then be applied to the 973 relevant local geostationary FRP product. For example, using a GOES-derived FRP product (Xu 974 et al., 2010) over eastern South America rather than the very high view zenith angle FRP data 975 provided by Meteosat SEVIRI. High latitudes prove more problematic for geostationary systems, 976 though Xu et al. (2010) did show that the high FRPs typical of crown fires in northern Canada 977 does, to some extent, offset increased northern latitude minimum FRP detection limits, and the 978 trade-off requires further investigation. With respect to (iii) and (iv), an increased number of fire 979 matchups could address both, by allowing for enough samples to stratify smoke emissions 980 coefficients by VZA as well as by biome. Including multiple years of data and focusing on 981 improving automation of the fire match-up process, for example through use of the MAIAC QA 982 product's 'smoke mask' (Lyapustin et al., 2012) and machine learning techniques, will help 983 increase the range and efficiency of the matchup process and thus the number of matchups used.

984

985 We compared our FREMv2 African biomass burning emission inventory data to those of GFAS, 986 GFED and two versions of FEER (FEER-GFASv1.2 and FEER-SEVIRI). FREMv2 provides the 987 highest levels of spatio-temporal detail (~0.05° spatially, updated 4 times per hour) since it can 988 exploit the native geostationary FRP data resolutions. Monthly FREMv2 TPM emission totals 989 agreed well with both FEER-GFASv1.2 and FEER-SEVIRI, and are significantly higher than those 990 of GFEDv4.1s and GFASv1.2 which past studies have suggested tend to underestimate NHAf 991 and SHAf aerosol emissions (Kaiser et al., 2012; Tosca et al., 2013; Ichoku and Ellison, 2014; 992 Reddington et al., 2016, Chevallier et al., 2009; Kopacz et al., 2010; Pechony et al., 2013). Trace 993 gas and carbon emissions are similarly also higher than for GFEDv4.1s and GFASv1.2, but similar

to those of FEER. Recent development of a 20 m African burned area (BA) product
(FireCCISFD11; Roteta et al., 2019) has shown a potential reason for this, because the MODIS
MCD64A1 500 m BA product, upon which GFED fire emissions estimates are based,
underestimates BA by as much as 50% in some regions of Africa, and GFAS is also indirectly
dependant on this BA product through its calibration against GFED (Kaiser et al., 2012).

999

Using the FREMv2 carbon emissions estimates we derive estimates of dry matter consumed (DMC) through an inversion of the Seiler and Crutzen (1980) approach, and dividing these by the FireCCISFD11 20 m BA product we deliver one of the first data-driven mappings of fuel consumption per unit area ( $F_c$ ) across Africa, which we find produces higher  $F_c$  in many areas compared to the modelled-based GFEDv4.1s (Figure 11).

1005

1006 Future developments advancing the FREM approach further will include its application to FRP 1007 data from other geostationary satellites based on the same baseline algorithm applied to generate 1008 the Meteosat FRP-PIXEL product used herein (Roberts et al., 2015; Wooster et al., 2015), for 1009 example data from Himawari (Xu et al., 2017), Meteosat Indian Ocean and GOES (Xu et al., 1010 2010). Since direct validation of large-scale fire emissions estimates remains challenging, future 1011 developments will also use the final FREMv2 smoke emissions estimates within atmospheric 1012 models to generate trace gas concentration and AOD fields for comparison to ground-based and 1013 EO-derived measures, exploiting a validation strategy similar to that previously used for 1014 evaluating other large scale fire emissions estimates (e.g. Baldassarre et al., 2015).

1015

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 53

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- 1030
- 1031
- 1032 *7. APPENDIX*
- **Appendix A -** CCI landcover classes and VCF Tree cover % assigned to each FREMv2 biome.

FREMv2 biome	Assigned CCI class codes	Landsat VCF Tree Cover above 5 m
closed canopy forest	50,180,160,170	-
low-woodland savanna	60,61,62,70,90,100,110	< 20 %
high-woodland savanna	60,61,62,70,90,100,110	> 20 %
shrubland	120,122,121	-
grassland	130,150,151,152,153	-
managed lands	10,11,12,20,30,40,190	-

1034

Appendix B - Mean percentage contribution of different biomes to the fire radiative energy (FRE) released by fires between 2013 and 2018. On average woodland savanna fires contribute the greatest total FRE throughout the year, except for in November, and this means their determination is especially important for overall smoke emission estimate accuracy. The importance and abundance of fires in this biome is reflected in the high numbers of fire-plume matchups identified for both the *low- and high- woodland savanna* biomes (Figure 5).

1042



1043

1045Appendix C – Percentage difference in fuel consumption per unit area ( $F_c$ , kg.m<sup>-2</sup>)1046calculated at a 0.25° resolution for 2016 African fires by FREMv2 (Figure 11a) and1047GFEDv4.1s (Figure 11b). The former provides significantly higher values in around half1048of the 0.25° grid cells. Figure 11c provides the per-biome  $F_c$  statistics coming from1049FREMv2.



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1429	9. LIST OF FIGURE CAPTIONS
1430	
1431	Figure 1.
1432	Example plume from a fire burning north of the Save River (Mozambique), imaged on the
1433	morning of $8^{th}$ October 2015 at 11:15 UTC via the Aqua satellites MODIS sensor (at a VZA
1434	40.6°). (a) 10 km MxD04_DB AOD product; (b) MxD04_DB field showing the number of 1 km
1435	reflectance pixels (out of 100) used to retrieve each 10 km AOD pixel value; (c) 500 m MODIS
1436	Corrected Reflectance (True Colour) image; and (d) 1 km MCD19 MAIAC AOD product derived
1437	from the same MODIS imagery shown in (c). The colour scale shown in (d) is also relevant for
1438	(a). The plume is far more easily distinguished in the 1 km than the 10 km AOD product and

- better matched to the smoke spatial distribution shown in the MODIS true colour image of (c).
- 1440 Unlike the 1 km AOD product of (d), the 10 km MxD04\_BD AOD data of (a) rather poorly
  - 73

defines the plume bounds and some pixels in this product are heavily impacted by the cloud
mask which removes AOD pixels over the thickest smoke (b). Some erroneous masking does
occur in the 1 km product of (d), shown as the black pixels, but this is minimal and addressed
via the interpolation described in Section 4.3.

1445

1446 **Figure 2.** 

1447 (a) Estimated fire emitted Total Particulate Matter (TPM) in 635 individual smoke plumes, as

derived from the 10 km MxD04\_DB AOD product (*orange*) and the 1 km MAIAC AOD product

1449 (*blue*), shown as a function of sensor view zenith angle (VZA). (b) Direct comparison of the

1450 matching MxD04\_DB and MAIAC-derived TPM values for each plume, restricted to plumes

1451 observed at VZA  $\leq$  20°. TPM is calculated from AOD using the equations presented in Ichoku

and Kaufman (2005) as described in Section 4.4.

1453

## 1454 **Figure 3**.

(a) Mapped percentage tree cover above 5 metres, as determined from the 30 m spatial resolution
Landsat Vegetation Continuous fields (VCF) product for 2015. (b) FREMv2 biome map for Africa
derived from the 2015 ESA CCI Landcover map (itself derived from 300 m PROBA-V imagery)
and the Landsat VCF product. Biomes were aggregated from the 36 land cover types defined in
the original CCI map, with the two woodland savanna biomes separated using (a) (see Appendix
A).

1461

1462 **Figure 4**.

1463 Example region of interest (ROI) containing the fire shown in Figure 1 along with matchup active 1464 fire pixels (white triangles) from the Meteosat FRP-PIXEL product detected between midnight 1465 and the MODIS overpass time. (a) MAIAC 1km AOD and SEVIRI active fire (AF) pixels; (b) 1466 histogram thresholding to discriminate the plume from the surrounding 'ambient' background. 1467 (c) SEVIRI AF pixels detected within the convex hull of the plume are considered to come from 1468 the same 'fire' that produced the smoke plume. The fire radiative energy (FRE) of the causal fire 1469 is then calculated from these observations and used to match to the AOD-derived total 1470 particulate matter (TPM) (see Figure 5).

1471

## 1472 **Figure 5.**

Smoke emission coefficients (C<sub>biome</sub>; in g.MJ<sup>-1</sup>) for the six African fire-affected biomes defined in 1473 1474 Section 4.1, each derived from the slope of an orthogonal distance regression (ODR) between 1475 data on fire-emitted total particulate matter (TPM) and matching fire radiative energy (FRE). Grev 1476 shaded area defines the 95% probability prediction interval of the ODR-derived slope. Each 1477 scatterplot is accompanied by an illustrative insert that depicts the typical landcover for the biome 1478 as seen in Google Earth (example locations are Closed Canopy Forest, 10.359° S, 19.086° E; 1479 Grassland 21.180° S, 19.560° E; Managed Land 10.495° N, 7.586° E; Low-Woodland Savanna 1480 7.085° N, 27.095° E, High-Woodland Savanna 12.523° S, 23.323° E; Shrubland 23.055° N, 1481 22.242° E).

- 1482
- 1483 **Figure 6.**
- 1484 (a) FREMv2 smoke emission coefficient ( $C_e$ ) mapped at 0.05°; (b) the matching 1° FEER  $C_e$

1485 product; and (c) the difference. In (c), FEER grid cells whose value was derived from gap-filling

1486 or which were calculated from less than 15 samples were removed (see Ichoku and Ellison 75 (2014) for full details). (d) Shows the spatial distribution of the fire matchups used to derive the
FREMv2 *C<sub>biome</sub>* values of Figure 5.

1489

1490 **Figure 7.** 

1491 Monthly total particulate matter (TPM) emissions from landscape fires for 2013 to 2018, as

derived using the FREMv2 methodology (blue) applied to the Meteosat FRP-PIXEL product of

1493 Wooster et al. (2015). Corresponding monthly TPM emissions are shown from GFEDv4.1s

1494 (green), GFASv1.2 (purple), and the FEERv1.0 coefficients applied to the GFASv1.2 (red) and

1495 SEVIRI FRE data (yellow).

1496

1497 **Figure 8.** 

1498Total particulate matter (TPM) emission density  $(g.m^{-2})$  across Africa for 2016 as determined by1499GFEDv4.1s (0.25° grid cells), GFASv1.2 (0.1° grid cells), FEERv1.0-GFASv1.2 (0.1° grid cells),1500FEERv1.0-SEVIRI (0.05° grid cells), and FREMv2 inventory derived herein (0.05° grid cells). The1501red 20° × 20° region outlined in the top left GFED plot is shown magnified for each inventory in1502the right-hand column.

1503 **Figure 9.** 

Mean monthly contribution (%) of fires to total particulate matter (TPM) emissions in each of the six FREMv2 biomes from 2013 to 2018 in (a) Northern Hemisphere Africa and (b) Southern Hemisphere Africa.

1507

1508	Figure	10.

Monthly total emissions (Tg) of (a) CO<sub>2</sub>; (b) CO; and (c) CH<sub>4</sub> for African landscape fires as estimated by FREMv2 (*blue*) between 2013 and 2018. Corresponding values from GFEDv4.1s (*green*), GFASv1.2 (*purple*), FEERv1.0-GFASv1.2 (*red*) and FEERv1.0-SEVIRI (*yellow*) are shown for comparison.

1513

1514 **Figure 11**.

1515 Fuel consumption per unit area ( $F_c$ , kg.m<sup>-2</sup>) mapped at 0.25° from (a) 2016 FREMv2 dry matter

1516 consumed (DMC) totals and the FireCCISFD11-estimated burned area, and (b) GFEDv4.1s. (c)

1517 Per-biome FREMv2 *F<sub>c</sub>* frequency distributions and derived means, medians and standard

1518 deviations. Note that  $F_c$  values in (a) apply only to the burned area patch inside a given pixel,

1519 and not the 0.25° pixel as a whole.

1520

1521

1522

Appendix A - CCI landcover classes and VCF Tree cover % assigned to each FREMv2 biome.
1524

**Appendix B** - Mean percentage contribution of different biomes to the fire radiative energy (FRE) released by fires between 2013 and 2018. On average woodland savanna fires contribute the greatest total FRE throughout the year, except for in November, and this means their determination is especially important for overall smoke emission estimate accuracy. The importance and abundance of fires in this biome is reflected in the high numbers of fire-plume matchups identified for both the *low- and high- woodland savanna* biomes (Figure 5).

1531

- **Appendix C** Percentage difference in fuel consumption per unit area ( $F_c$ , kg.m<sup>-2</sup>)
- 1533 calculated at a 0.25° resolution for 2016 African fires by FREMv2 (Figure 11a) and
- 1534 GFEDv4.1s (Figure 11b). The former provides significantly higher values in around half
- 1535 of the 0.25° grid cells. Figure 11c provides the per-biome  $F_c$  statistics coming from
- 1536 FREMv2.